

# Spherical Geometry

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## 1 Spherical coordinates

In order to simplify notation, we define the operators

$$\begin{aligned} \mathbf{d}_r &= \partial_r & (1) \\ \mathbf{d}_\theta &= r^{-1} \partial_\theta & (2) \\ \mathbf{d}_\phi &= \varpi^{-1} \partial_\phi & (3) \end{aligned}$$

where  $\varpi = r \sin \theta$  is the cylindrical radius, and write the gradient as

$$\text{grad } \Psi = \begin{pmatrix} \mathbf{d}_r \Psi \\ \mathbf{d}_\theta \Psi \\ \mathbf{d}_\phi \Psi \end{pmatrix} \quad (4)$$

The divergence and curl operators are written as

$$\text{div } \mathbf{A} = \mathcal{D}_r A_r + \mathbf{D}_\theta A_\theta + \mathbf{d}_\phi A_\phi \quad (5)$$

and

$$\text{curl } \mathbf{A} = \begin{pmatrix} \mathbf{D}_\theta A_\phi - \mathbf{d}_\phi A_\theta \\ \mathbf{d}_\phi A_r - \mathbf{D}_r A_\phi \\ \mathbf{D}_r A_\theta - \mathbf{d}_\theta A_r \end{pmatrix}, \quad (6)$$

and the laplacian of a scalar,  $\nabla^2 \equiv \text{div grad}$ , is

$$\nabla^2 \Psi = \mathbf{D}_r^2 \Psi + \mathbf{D}_\theta^2 \Psi + \mathbf{d}_\phi^2 \Psi; \quad (7)$$

see Table 1 for the definition of additional operators, and Table 2 for the relations between them.

Table 1: Useful operators in spherical coordinates.

$\mathcal{D}_r = r^{-2} \partial_r (r^2 \cdot)$	
$\mathbf{D}_r = r^{-1} \partial_r (r \cdot)$	$\mathbf{D}_\theta = r^{-1} \sin^{-1} \theta \partial_\theta (\sin \theta \cdot)$
$\mathbf{D}_r^{-1} = r \partial_r (r^{-1} \cdot)$	$\mathbf{D}_\theta^{-1} = r^{-1} \sin \theta \partial_\theta (\sin^{-1} \theta \cdot)$

Express all spatial derivatives in fully non-conservative form. Arrange mixed operators such that the innermost

Table 2: Some useful identities. The ‘ $\equiv$ ’ symbol indicates the definition of an operator, so  $\mathcal{D}_r^2$  is not meant to be the same as the operator  $\mathcal{D}_r$  applied twice.

$\mathbf{d}_r^2 = \mathbf{d}_r \mathbf{d}_r$	$\mathbf{d}_\theta^2 = \mathbf{d}_\theta \mathbf{d}_\theta$	$\mathbf{d}_\phi^2 = \mathbf{d}_\phi \mathbf{d}_\phi$
$\mathbf{D}_r^2 \equiv \mathbf{D}_r \mathbf{D}_r = \mathcal{D}_r \mathbf{d}_r$	$\mathbf{D}_\theta^2 \equiv \mathbf{D}_\theta \mathbf{d}_\theta$	
$\mathcal{D}_r^2 \equiv \mathbf{d}_r \mathcal{D}_r = \mathbf{D}_r^{-1} \mathbf{d}_r$	$\mathcal{D}_\theta^2 \equiv \mathbf{d}_\theta \mathbf{D}_\theta$	
	$\mathbf{d}_r \mathbf{D}_\theta = \mathbf{D}_\theta \mathbf{D}_r^{-1}$	$\mathbf{d}_\theta \mathbf{d}_\phi = \mathbf{d}_\phi \mathbf{D}_\theta^{-1}$

derivative operator is in the  $r$ -direction, the next one in the  $\theta$ -direction, and the  $\phi$ -operator is the outermost one.

$$\text{curl}^2 \mathbf{A} = \begin{pmatrix} \mathbf{D}_\theta \mathbf{D}_r A_\theta + \mathbf{d}_\phi \mathbf{D}_r A_\phi - \mathbf{D}_\theta^2 A_r - \mathbf{d}_\phi^2 A_r \\ \mathbf{d}_\phi \mathbf{D}_\theta A_\phi + \mathbf{d}_\theta \mathbf{D}_r A_r - \mathbf{d}_\phi^2 A_\theta - \mathbf{D}_r^2 A_\theta \\ \mathbf{d}_\phi \mathbf{D}_r A_r + \mathbf{d}_\theta \mathbf{d}_\phi A_\theta - \mathbf{D}_r^2 A_\phi - \mathbf{D}_\theta^2 A_\phi \end{pmatrix} \quad (8)$$

is the double-curl.

$$\text{grad div } \mathbf{A} = \begin{pmatrix} \mathbf{d}_r \mathcal{D}_r A_r + \mathbf{d}_r \mathbf{D}_\theta A_\theta + \mathbf{d}_r \mathbf{d}_\phi A_\phi \\ \mathbf{d}_\theta \mathcal{D}_r A_r + \mathbf{d}_\theta \mathbf{D}_\theta A_\theta + \mathbf{d}_\theta \mathbf{d}_\phi A_\phi \\ \mathbf{d}_\phi \mathcal{D}_r A_r + \mathbf{d}_\phi \mathbf{D}_\theta A_\theta + \mathbf{d}_\phi \mathbf{d}_\phi A_\phi \end{pmatrix} \quad (9)$$

or, written such that  $r$ -operators come first and  $\phi$  operators last,

$$\text{grad div } \mathbf{A} = \begin{pmatrix} \mathcal{D}_r^2 A_r + \mathbf{D}_\theta \mathbf{D}_r^{-1} A_\theta + \mathbf{d}_\phi \mathbf{D}_r^{-1} A_\phi \\ \mathbf{d}_\theta \mathcal{D}_r A_r + \mathcal{D}_\theta^2 A_\theta + \mathbf{d}_\phi \mathbf{D}_\theta^{-1} A_\phi \\ \mathbf{d}_\phi \mathcal{D}_r A_r + \mathbf{d}_\phi \mathbf{D}_\theta A_\theta + \mathbf{d}_\phi^2 A_\phi \end{pmatrix} \quad (10)$$

The laplacian of a vector is evaluated as  $\nabla^2 = \text{grad div} - \text{curl curl}$  and can be written as

$$\nabla^2 \mathbf{A} = \begin{pmatrix} (\nabla^2 - 2r^{-2}) A_r - 2r^{-1} (\mathbf{D}_\theta A_\theta + \mathbf{d}_\phi A_\phi) \\ (\nabla^2 - \varpi^{-2}) A_\theta + 2r^{-1} (\mathbf{d}_\theta A_r - \cot \theta \mathbf{d}_\phi A_\phi) \\ (\nabla^2 - \varpi^{-2}) A_\phi + 2r^{-1} (\mathbf{d}_\phi A_r + \cot \theta \mathbf{d}_\phi A_\theta) \end{pmatrix} \quad (11)$$

Table 3: Spatial operators relevant in spherical coordinates.  $\varpi = r \sin \theta$  is the cylindrical radius.

$d_r = \partial_r$	$d_\theta = r^{-1} \partial_\theta$	$d_\phi = \varpi^{-1} \partial_\phi$
$D_r = r^{-1} + d_r$	$D_\theta = r^{-1} \cot \theta + d_\theta$	
$D_r^{-1} = -r^{-1} + d_r$	$D_\theta^{-1} = -r^{-1} \cot \theta + d_\theta$	
$\mathcal{D}_r = 2r^{-1} + d_r$		
$d_r^2 = \partial_r^2$	$d_\theta^2 = r^{-2} \partial_\theta^2$	$d_\phi^2 = \varpi^{-2} \partial_\phi^2$
$D_r^2 = 2r^{-1} d_r + d_r^2$	$D_\theta^2 = r^{-1} \cot \theta d_\theta + d_\theta^2$	
$\mathcal{D}_r^2 = -2r^{-2} + D_r^2$	$\mathcal{D}_\theta^2 = -\varpi^{-2} + D_\theta^2$	

## 2 ... without radial derivatives

$$d_\phi d_\phi = \varpi^{-2} \partial_{\phi\phi} \quad (23)$$

Useful for solar surface observations

curl2

$$\text{curl } \mathbf{A} = \begin{pmatrix} D_\theta A_\phi - d_\phi A_\theta \\ +d_\phi A_r \\ -d_\theta A_r \end{pmatrix}, \quad (12)$$

$$D_r^2 = 2r^{-1} \partial_r + \partial_{rr} \quad (24)$$

Compute  $A_r \equiv T \hat{r}$  from  $J_r = L^2 A_r$ .

$$\text{grad div } \mathbf{A} = (\text{grad div } \mathbf{A})_{\text{cart}} + (\text{grad div } \mathbf{A})_{\text{extr}} \quad (25)$$

$$\text{curl}^2 \mathbf{A} = \begin{pmatrix} -D_\theta^2 A_r - d_\phi^2 A_r \\ +d_\phi D_\theta A_\phi - d_\phi^2 A_\theta \\ +d_\theta d_\phi A_\theta - \mathcal{D}_\theta^2 A_\phi \end{pmatrix} \quad (13)$$

$$(\text{grad div } \mathbf{A})_{\text{cart}} = \quad (26)$$

$$\begin{pmatrix} A_{r,rr} + r^{-1} A_{\theta,\theta r} + \varpi^{-1} A_{\phi,\phi r} \\ r^{-1} A_{r,r\theta} + r^{-2} A_{\theta,\theta\theta} + r^{-1} \varpi^{-1} A_{\phi,\phi\theta} \\ \varpi^{-1} A_{r,r\phi} + r^{-1} \varpi^{-1} A_{\theta,\theta\phi} + \varpi^{-2} A_{\phi,\phi\phi} \end{pmatrix} \quad (27)$$

## 3 Modified cartesian operators

Consider first

$$\text{grad div } \mathbf{A} = \begin{pmatrix} d_r \mathcal{D}_r A_r + d_r D_\theta A_\theta + d_r d_\phi A_\phi \\ d_\theta \mathcal{D}_r A_r + d_\theta D_\theta A_\theta + d_\theta d_\phi A_\phi \\ d_\phi \mathcal{D}_r A_r + d_\phi D_\theta A_\theta + d_\phi d_\phi A_\phi \end{pmatrix} \quad (14)$$

## 4 Legendre polynomials

$$A_\theta = d_\phi S, \quad A_\phi = -d_\theta S. \quad (29)$$

Write explicitly

Backward transformation

$$d_r \mathcal{D}_r = -2r^{-2} + 2r^{-1} \partial_r + \partial_{rr} \quad (15)$$

$$d_r D_\theta = -r^{-2} \cot \theta + r^{-1} \cot \theta \partial_r - r^{-2} \partial_\theta + r^{-1} \partial_{r\theta} \quad (16)$$

$$d_r d_\phi = -r^{-1} \varpi^{-1} \partial_\phi + \varpi^{-1} \partial_{r\phi} \quad (17)$$

$$d_\theta \mathcal{D}_r = 2r^{-2} \partial_\theta + r^{-1} \partial_{r\theta} \quad (18)$$

$$d_\theta D_\theta = -r^{-2} (1 + \cot^2 \theta) + r^{-2} \cot \theta \partial_\theta + r^{-2} \partial_{\theta\theta} \quad (19)$$

$$d_\theta d_\phi = -r^{-1} \varpi^{-1} \cot \theta \partial_\phi + r^{-1} \varpi^{-1} \partial_{\theta\phi} \quad (20)$$

$$d_\phi D_r = r^{-1} \varpi^{-1} \partial_\phi + \varpi^{-1} \partial_{r\phi} \quad (21)$$

$$d_\phi D_\theta = r^{-1} \varpi^{-1} \cot \theta \partial_\phi + r^{-1} \varpi^{-1} \partial_{\theta\phi} \quad (22)$$

$$S_m(r, \theta) = \sum_{\ell=1}^L \sqrt{N_{\ell m}} S_{\ell m}(r) P_\ell^m(\cos \theta), \quad (30)$$

where

$$N_{\ell m} = \frac{1}{2} (2\ell + 1) \frac{(\ell - m)!}{(\ell + m)!} \quad (31)$$

are the normalization coefficients of the Legendre polynomials. Forward transformation

$$S_{\ell m}(r) = \sqrt{N_{\ell m}} \int_0^\pi P_\ell^m(\cos \theta) S(r, \theta) \sin \theta d\theta \quad (32)$$

Useful relation

$$\sin \theta \frac{dP_\ell^m}{d\theta} = -\frac{(\ell+1)(\ell+m)}{2\ell+1} P_{\ell-1}^m + \frac{\ell(\ell-m+1)}{2\ell+1} P_{\ell+1}^m \quad (33)$$

$$\cos \theta P_\ell^m = \frac{\ell+m}{2\ell+1} P_{\ell-1}^m + \frac{\ell-m+1}{2\ell+1} P_{\ell+1}^m \quad (34)$$

## 5 Spherical Bessel functions

$$j_0(z) = \frac{\sin z}{z} \equiv \sqrt{\pi/2z} J_{1/2}(z) \quad (35)$$

$$j_1(z) = \frac{\sin z}{z^2} - \frac{\cos z}{z} \quad (36)$$

$$j_2(z) = \left( \frac{3}{z^3} - \frac{1}{z} \right) \sin z - \frac{3}{z} \cos z \quad (37)$$

Zeros:  $j_l(z_l) = 0$ ,  $z_0 = \pi$ ,  $z_1 = 4.493409$ ,  $z_2 = 5.763459$ .  
Note that

$$dj_1/d \ln z|_{z_1} = -1. \quad (38)$$

## 6 Relative magnetic helicity

$$H_{\text{rel}} = \int (\mathbf{A} + \mathbf{A}_P) \cdot (\mathbf{B} - \mathbf{B}_P) dV, \quad (39)$$

where  $\mathbf{B}_P = \nabla \times \mathbf{A}_P$  is a potential (current-free) reference field in the interior with  $\mathbf{B}_{Pr} = \mathbf{B}_r$  on  $r = R$ . We express  $\mathbf{A}_P$  in terms of the poloidal part of the superpotential,

$$\mathbf{A}_P = \nabla \times (\hat{\mathbf{r}}\Phi). \quad (40)$$

The corresponding current is

$$\mathbf{J}_P = \nabla \times \nabla \times \nabla \times (\hat{\mathbf{r}}\Phi) = -\nabla \times (\hat{\mathbf{r}}\nabla^2\Phi). \quad (41)$$

The potential field is current-free, i.e.  $\mathbf{J}_P = 0$ , so we require

$$\nabla^2\Phi = 0. \quad (42)$$

We solve this equation in spectral space and define

$$\Phi_\ell(r) = \sqrt{N_{\ell m}} \int_0^\pi P_\ell(\cos \theta) \Phi(r, \theta) \sin \theta d\theta. \quad (43)$$

where  $m = 0$ . The backward transform is

$$\Phi(r, \theta) = \sum_{\ell=1}^L \sqrt{N_{\ell m}} \Phi_\ell(r) P_\ell(\cos \theta), \quad (44)$$

and

$$N_{\ell m} = \frac{1}{2}(2\ell+1) \frac{(\ell-m)!}{(\ell+m)!} \quad (45)$$

are the normalization coefficients of the Legendre polynomials.

It is convenient to define  $\tilde{\Phi}_\ell = r\Phi_\ell$ , so the potential equation in spectral space reads

$$\tilde{\Phi}_\ell'' - \frac{\ell(\ell+1)}{r^2} \tilde{\Phi}_\ell = 0, \quad (46)$$

where and primes denote radial derivatives. Since the radial components of  $\mathbf{B}$  and  $\mathbf{B}_P$  coincide on  $r = R$ , we have

$$B_r = -\mathbf{r} \cdot \nabla \times \nabla \times \mathbf{r}\Phi = -\nabla_\perp^2 \tilde{\Phi}. \quad (47)$$

on  $r = R$ . In spectral space this corresponds to the boundary condition

$$\tilde{\Phi}_\ell(R) = \frac{R^2}{\ell(\ell+1)} b_\ell \quad (48)$$

where

$$b_\ell = \sqrt{N_{\ell m}} \int_0^\pi P_\ell(\cos \theta) B_r(R, \theta) \sin \theta d\theta. \quad (49)$$

## 7 Useful relations in axisymmetry

A useful relation in connection with axisymmetry is

$$\mathcal{D}_\theta^2 d_\theta = d_\theta D_\theta d_\theta = d_\theta D_\theta^2. \quad (50)$$

It allows to show that the vacuum condition in axisymmetry,  $D^2 A_\phi = 0$ , is equivalent to the general vacuum condition,  $\nabla^2 \Phi = 0$ , where  $\Phi$  is the poloidal superpotential of the vacuum field, so  $\mathbf{A} = \nabla \times (\hat{\mathbf{r}}\Phi)$ . Here we have introduced the operator

$$D^2 = -\varpi^{-2} + \nabla^2 = D_r^2 + \mathcal{D}_\theta^2, \quad (51)$$

which is frequently used in axisymmetry. In axisymmetry, only the  $\phi$  component of the  $\mathbf{A}$  of a potential field is finite, and  $A_\phi = r d_\theta \Phi$ . Thus,

$$D^2 A_\phi = D^2 r d_\theta \Phi = r d_\theta D_r^2 \Phi + r d_\theta \mathcal{D}_\theta^2 \Phi = r d_\theta \nabla^2 \Phi = 0 \quad (52)$$

Another useful relation in axisymmetry is

$$\frac{dP_\ell(\cos \theta)}{d\theta} = P_\ell^1(\cos \theta) \quad (53)$$

## 7.1 Examples

$$D_r^2 j_1(kr) = (2r^{-2} - k^2) j_1(kr) \quad (54)$$

$$D_\theta^2 \sin \theta = -2r^{-2} \sin \theta \quad (55)$$

Thus

$$b(r, \theta) = j_1(kr) \sin \theta \quad (56)$$

is an eigenfunction of  $D^2$  with

$$D^2 b + k^2 b = 0. \quad (57)$$

A particular solution of the potential equation

$$D^2 a = 0 \quad (58)$$

is

$$a = \begin{cases} \varpi & \text{for } r \leq 1, \\ \varpi/r^3 & \text{for } r > 1. \end{cases} \quad (59)$$

where  $\varpi = r \sin \theta$ .

## 7.2 Constant $\alpha$ sphere

The axisymmetric steady state dynamo equations for  $\alpha = \text{const}$  and  $\eta_T = R = 1$  are

$$\alpha b + D^2 a = 0, \quad -\alpha D^2 a + D^2 b = 0. \quad (60)$$

A solution that satisfies the boundary condition

$$da/dr + 2a = 0 \quad \text{on } r = 1 \quad (61)$$

is

$$a = k^{-1} [j_1(kr) + \frac{1}{3}r] \sin \theta, \quad b = j_1(kr) \sin \theta \quad (62)$$

(with  $k = 4.493409$ ) in  $r \leq 1$  and

$$a = \frac{1}{3}k^{-1}r^{-2} \sin \theta, \quad b = 0 \quad (63)$$

in  $r > 1$ .

## 7.3 Rewrite

$$\mathbf{A} = a\hat{\phi} + \nabla \times c\hat{\phi}, \quad (64)$$

$$\mathbf{B} = b\hat{\phi} + \nabla \times a\hat{\phi}, \quad (65)$$

where  $D^2 c = -b$  and  $c = 0$  on  $r = 1$  is *assumed*. Therefore,

$$c = k^{-2} j_1(kr) \sin \theta \quad (66)$$

## 7.4 The potential part of $B$

Both in the interior and the exterior the field has a potential part,  $\nabla \Psi$ . This can be seed by splitting

$$a = a_0 + a_1 \quad (67)$$

with

$$a_0 = k^{-1} j_1(kr) \sin \theta, \quad a_1 = \frac{1}{3}k^{-1}r \sin \theta, \quad (68)$$

where  $a_1$  satisfies

$$\mathbf{B}_1 = \nabla \times a_1 \hat{\phi} = \nabla \Psi. \quad (69)$$

In this case,  $\mathbf{B}_1 = \frac{2}{3}k^{-1}\mathbf{z}$  (in  $r < 1$ ), so

$$\Psi = \frac{2}{3}k^{-1}r \cos \theta \quad r < 1, \quad (70)$$

and

$$\Psi = \frac{1}{3}k^{-1}r^{-2} \cos \theta \quad r > 1. \quad (71)$$

## 7.5 Vacuum reference field

$$\mathbf{A}_P = a_P \hat{\phi} + \nabla \times c_P \hat{\phi}, \quad (72)$$

$$\mathbf{B}_P = b_P \hat{\phi} + \nabla \times a_P \hat{\phi}, \quad (73)$$

$$a_P = \frac{1}{3}rk^{-1} \sin \theta, \quad b_P = 0, \quad c_P = 0. \quad (74)$$

## 7.6 Integral over P

$$\int_0^\pi P_\ell(\cos \theta) \sin \theta d\theta = \frac{2}{2\ell + 1} \quad (75)$$

## 8 Magnetic helicity in spheres embedded in a vacuum

The helicity integral can be split into toroidal and poloidal components, i.e.

$$\mathbf{A} \cdot \mathbf{B} = ab + \mathbf{A}_P \cdot \mathbf{B}_P \quad (76)$$

The poloidal part of the helicity integral can be written in the form

$$\mathbf{A}_P \cdot \mathbf{B}_P = ab + \nabla \cdot (a\hat{\phi} \times \mathbf{A}_P). \quad (77)$$

This shows first of all that the magnetic helicity for the infinite volume (surface integral vanishes) can be written as

$$H = \int_{V+V_{\text{ext}}} (ab + \mathbf{A}_p \cdot \mathbf{B}_p) dV = 2 \int_V ab dV. \quad (78)$$

This is gauge-invariant, because there is no surface term. Since  $b = 0$  outside the sphere, this integral is only over the sphere. Except for the case of the constant  $\alpha$  sphere,  $H$  will always be zero because of cancelation between the two hemispheres. Thus, the magnetic helicity integral for the infinite upper half space is, in the Coulomb gauge,

$$H_N^{(\text{Cou})} = 2 \int_N ab dV + \int_{\theta=\pi/2} a \mathbf{A}_p \cdot d\mathbf{S}. \quad (79)$$

The second integral vanishes for fields with dipolar symmetry ( $A_r = 0$  on the midplane).

## 8.1 Relative helicity

We make use of the fact that for the vacuum field only  $a_p$  is finite. The contribution from the toroidal components to the relative helicity integral is just  $\int (a + a_p) b dV$ , because the vacuum field does not have a toroidal field. For the poloidal contribution we note that the vacuum field has a purely toroidal vector potential, so we have

$$\mathbf{A}_p \cdot (\mathbf{B} - \mathbf{B}_p)_p = b(a - a_p) + \nabla \cdot [(a - a_p) \hat{\phi} \times \mathbf{A}_p]. \quad (80)$$

However, on the boundary we always have

$$a = a_p \quad \text{on } \partial V, \quad (81)$$

so the relative magnetic helicity is always (both for toroidal and poloidal fields)

$$H^{(\text{rel})} = 2 \int_V ab dV. \quad (82)$$

In particular, we have

$$H_N^{(\text{rel})} = 2 \int_N ab dV. \quad (83)$$

## 8.2 Mean-field $\alpha\Omega$ dynamo

We consider here the case without meridional circulation, so we have

$$(\partial_t - \eta_T D^2) a = \alpha b \quad (84)$$

$$(\partial_t - \eta_T D^2) b = \hat{\phi} \cdot \nabla \times \alpha \mathbf{B}_p + \varpi \mathbf{B}_p \cdot \nabla \Omega. \quad (85)$$

We are interested in the evolution of  $ab$ , so from the first equation we have

$$b \partial_t a = \alpha b^2 - \eta_T j b, \quad (86)$$

where  $j = -D^2 a$ . From the second equation we have

$$a \partial_t b = \alpha \mathbf{B}_p^2 - \eta_T \mathbf{J}_p \cdot \mathbf{B}_p - \frac{1}{2} \nabla \cdot \mathbf{F}_H. \quad (87)$$

The factor  $1/2$  enters in the expression for the helicity flux, because the relative magnetic helicity is given by  $2ab$ , so

$$\mathbf{F}_H = -2(\alpha \mathbf{B}_p - \eta_T \mathbf{J}_p) \times \hat{\phi} a - 2a \varpi \Omega \mathbf{B}_p, \quad (88)$$

where we have made use of the identities

$$a \hat{\phi} \cdot \nabla \times \alpha \mathbf{B}_p = \alpha \mathbf{B}_p^2 + \alpha \mathbf{B}_p \times \hat{\phi} a, \quad (89)$$

$$a D^2 b = -\mathbf{J}_p \cdot \mathbf{B}_p - \nabla \cdot \mathbf{J}_p \times \hat{\phi} a, \quad (90)$$

$$\varpi a \mathbf{B}_p \cdot \nabla \Omega = \nabla \cdot (a \varpi \Omega \mathbf{B}_p). \quad (91)$$

The last one comes from the helicity flux due to the toroidal velocity,  $\mathbf{U} = \varpi \Omega \hat{\phi}$ ,

$$(\mathbf{U} \times \mathbf{B}_p) \times a \hat{\phi} = a \varpi \Omega \mathbf{B}_p \quad (92)$$

We also note that in axisymmetry the poloidal magnetic field can be written in the form

$$\mathbf{B}_p \equiv \nabla \times a \hat{\phi} = -\varpi^{-1} \hat{\phi} \times \nabla (\varpi a) \quad (93)$$

Thus, the final magnetic helicity equation is

$$\frac{d}{dt} H^{(\text{rel})} = 2 \int_V \bar{\mathcal{E}} \cdot \bar{\mathbf{B}} - \oint_{\partial V} \mathbf{F}_H \cdot d\mathbf{S} \quad (94)$$

where

$$\bar{\mathcal{E}} \cdot \bar{\mathbf{B}} = \alpha \mathbf{B}^2 - \eta_T \mathbf{J} \cdot \mathbf{B} \quad (95)$$

where we have omitted the overbars again, and

$$\mathbf{F}_H = 2 \mathbf{E}_p \times \mathbf{A}_t, \quad (96)$$

where  $\mathbf{A}_t = \hat{\phi} a$  and

$$\mathbf{E}_p = -\mathbf{U} \times \mathbf{B}_p - (\alpha \mathbf{B}_p - \eta_T \mathbf{J}_p). \quad (97)$$

This equation is true even if there was meridional circulation, because the poloidal flow does not enter, so that

$$\mathbf{E}_p = -\mathbf{U}_t \times \mathbf{B}_p - (\alpha \mathbf{B}_p - \eta_T \mathbf{J}_p) \quad (98)$$

is the relevant expression even in the general case (although then the symbol  $\mathbf{E}_p$  is misleading, because one would think there would be the additional term  $\mathbf{U}_p \times \mathbf{B}_t$  which is not the case). On the other hand,  $\mathbf{U}_p = 0$  on the boundary, so there would not have been any confusion anyway. The expression (88) is therefore to be preferred.

### 8.3 Advective gauge

Assume  $\mathbf{U} = (0, 0, \varpi\Omega)$ , so

$$\frac{\partial \mathbf{A}}{\partial t} = \mathbf{U} \times \mathbf{B} = \Omega \nabla (\varpi A_\phi) = -\varpi A_\phi \nabla \Omega - \nabla \phi \quad (99)$$

with  $\phi = U_\phi A_\phi$ .