

# Batchelor, Saffman, and Kazantsev spectra in galactic small-scale dynamos

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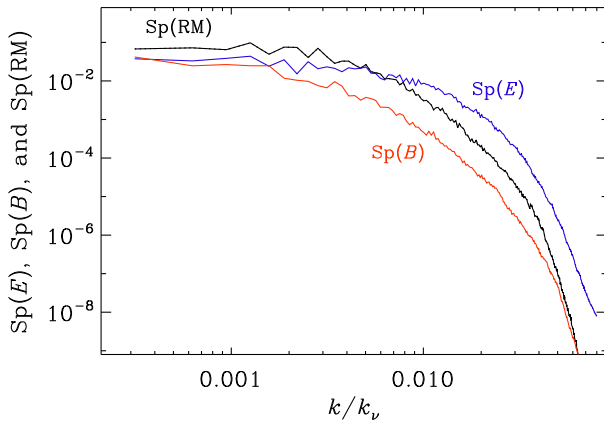
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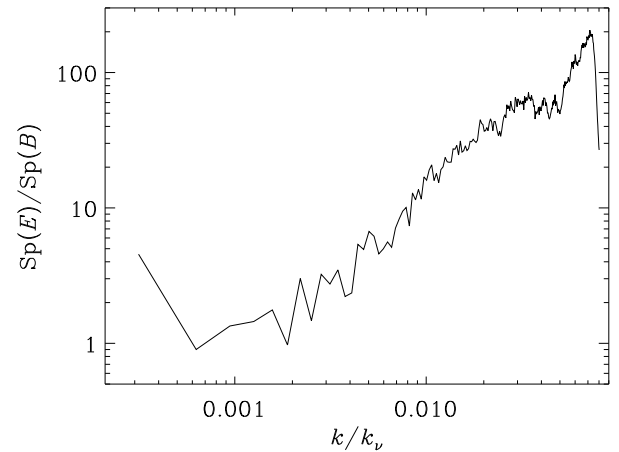
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## ABSTRACT

**Key words:** dynamo – MHD – polarization – turbulence – galaxies: magnetic fields



**Figure 1.** Same as Fig. ??, but for Run T0 during the kinematic stage.



**Figure 2.** Ratio

## ACKNOWLEDGEMENTS

### DATA AVAILABILITY

The source code used for the simulations of this study, the PENCIL CODE (Pencil Code Collaboration et al. 2021), is freely available on <https://github.com/pencil-code/>. The DOI of the code is <https://doi.org/10.5281/zenodo.2315093>. The simulation setups and the corresponding secondary data, as well as supplemental material with additional plots for the PDFs of Run B and diagnostic images and spectra for Runs B and D in the saturated state are available on <https://doi.org/10.5281/zenodo.7281479>; see also <https://www.nordita.org/~brandenb/projects/Kazantsev-Subinertial> for easier access to the same material as on the Zenodo site.

## REFERENCES

Pencil Code Collaboration et al., 2021, *JOSS*, **6**, 2807

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